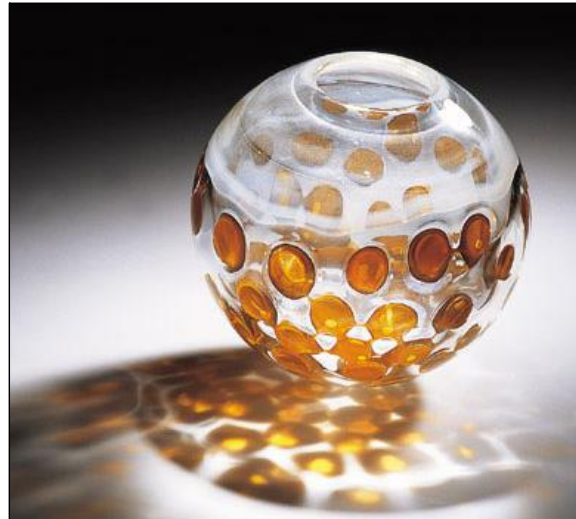
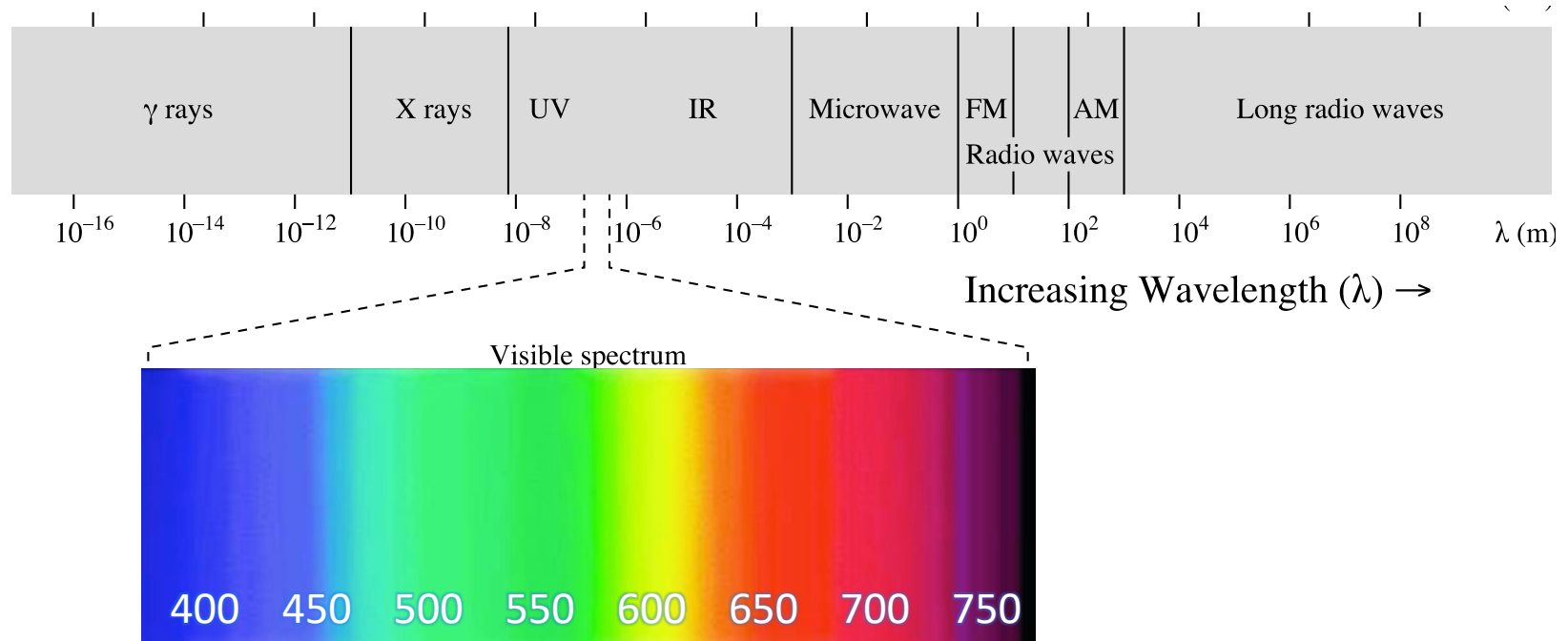


# Lighting and Shading I

Prof. Dr. Markus Gross

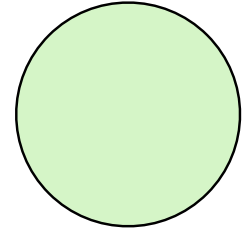


# Before: nature of light and colors

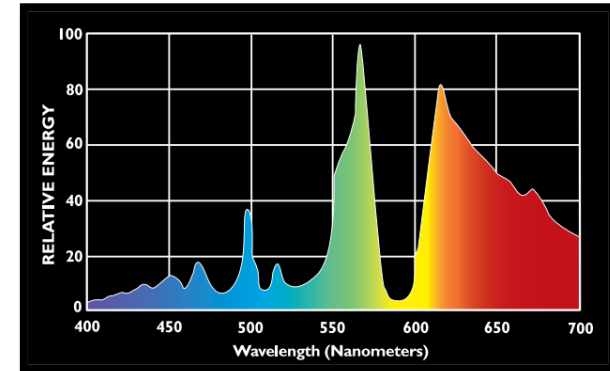


# Before: nature of light and colors

- Light can be a mixture of many wavelengths
- Spectral power distribution (SPD)
  - $P(\lambda)$  = intensity at wavelength  $\lambda$
  - intensity as a function of wavelength
- We perceive these distributions as colors

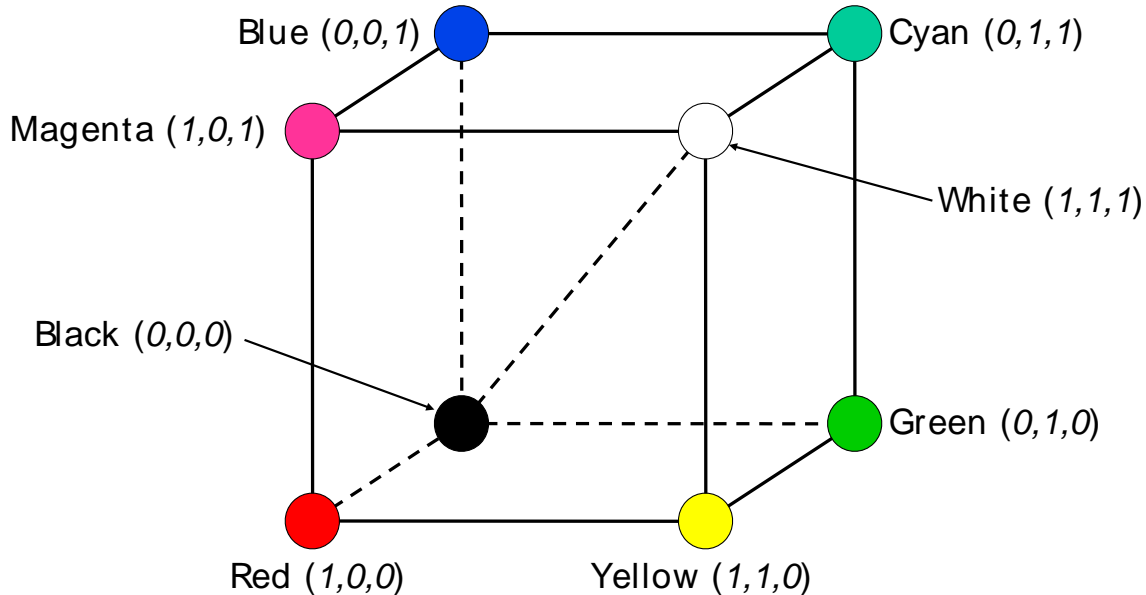


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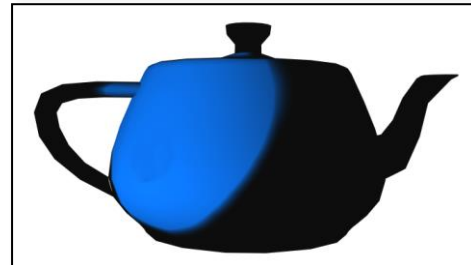
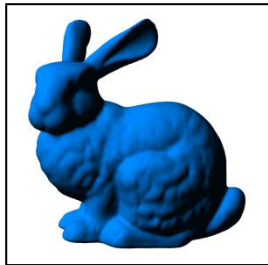
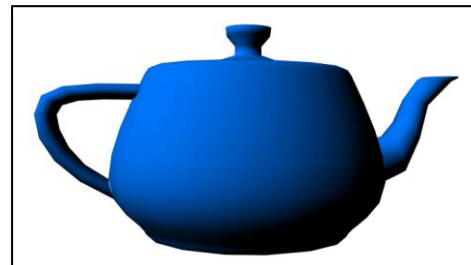
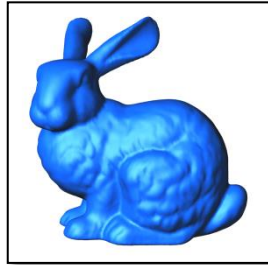
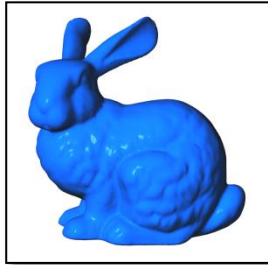
# Before: representing colors

- Unit cube with R,G,B basis vectors



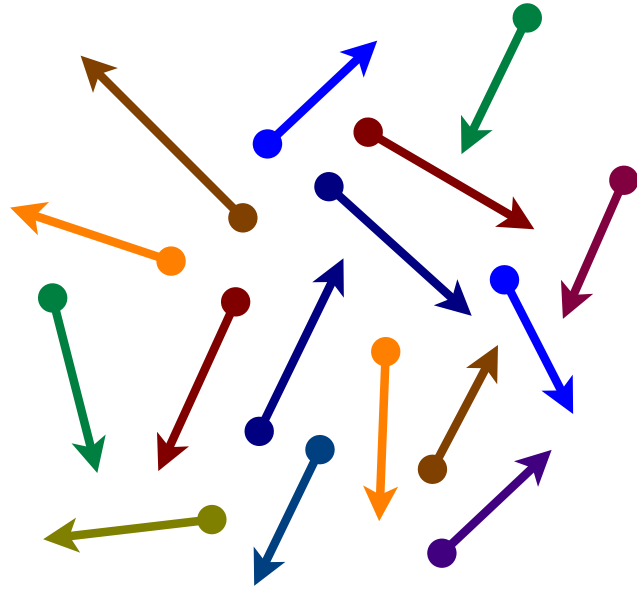
# Before: we need material models

- Interaction of light with geometry



# Measuring Light

- How do we measure light



Measuring = Counting photons

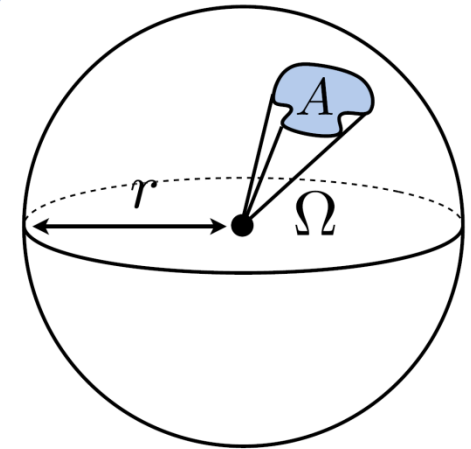
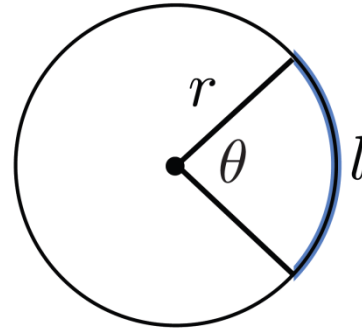
# Measuring Light

---

- Radiometry
  - Studies the measurement of electromagnetic radiation, including visible light

# Basic Definitions

- Angle:  $\theta = \frac{l}{r}$ 
  - circle:  $2\pi$  radians
- Solid angle:  $\Omega = \frac{A}{r^2}$ 
  - sphere:  $4\pi$  steradians





# Basic Definitions

- Direction

- point on the unit sphere
- parameterized by two angles

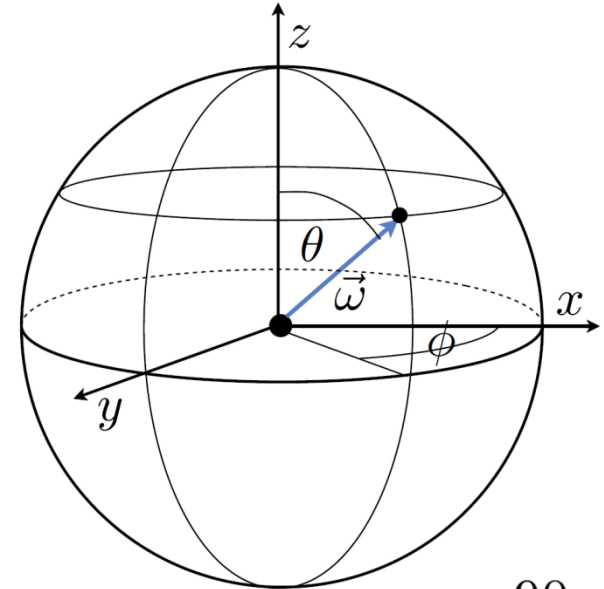
$$\vec{\omega}_x = \sin \theta \cos \phi$$

$$\vec{\omega}_y = \sin \theta \sin \phi$$

$$\vec{\omega}_z = \cos \theta$$

$$\vec{\omega} = (\theta, \phi)$$

↑                  ↑  
zenith          azimuth



$$\text{latitude} = \frac{90}{\pi}(\pi - \theta)$$

$$\text{longitude} = \frac{90}{\pi}\phi$$

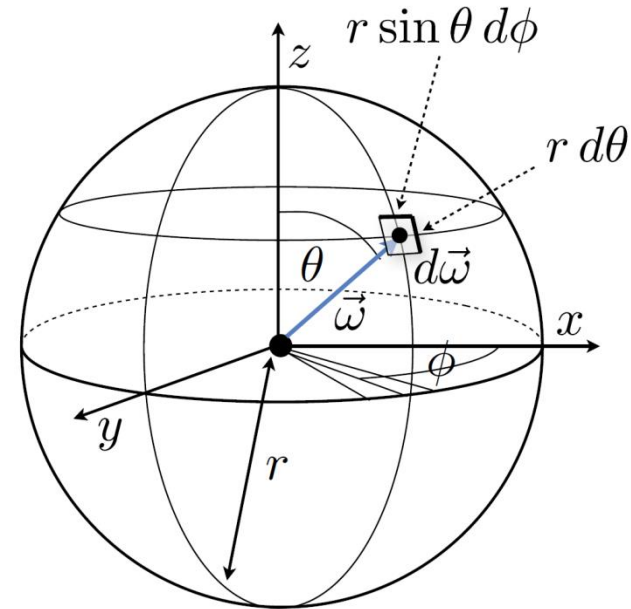
# Basic Definitions

- Differential Solid Angle

$$dA = (r d\theta)(r \sin \theta d\phi)$$

$$d\vec{\omega} = \frac{dA}{r^2} = \sin \theta d\theta d\phi$$

$$\Omega = \int_{S^2} d\vec{\omega} = \int_0^{2\pi} \int_0^\pi \sin \theta d\theta d\phi = 4\pi$$



# Basic Definitions

- Assume light consists of photons with
  - $\mathbf{x}$  : Position
  - $\vec{\omega}$  : Direction of motion
  - $\lambda$  : Wavelength
- Each photon has an energy of:  $\frac{hc}{\lambda}$ 
  - $h \approx 6.63 \cdot 10^{-34} \text{ m}^2 \cdot \text{kg}/\text{s}$  : Planck's constant
  - $c = 299,792,458 \text{ m}/\text{s}$  : speed of light in vacuum
  - Unit of energy, Joule :  $[J = \text{kg} \cdot \text{m}^2/\text{s}^2]$

# Radiometry

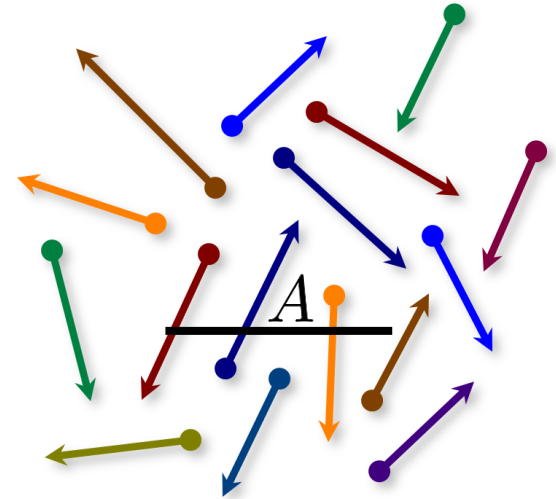
- Basic quantities
  - flux  $\Phi$
  - irradiance  $E$
  - radiosity  $B$
  - intensity  $I$
  - radiance  $L$

# Radiometry

- Flux (radiant flux, power)
  - total amount of energy passing through surface or space per unit time

$$\Phi(A) \quad \left[ \frac{J}{s} = W \right]$$

- examples:
  - number of photons hitting a wall per second
  - number of photons leaving a lightbulb per second



# Radiometry

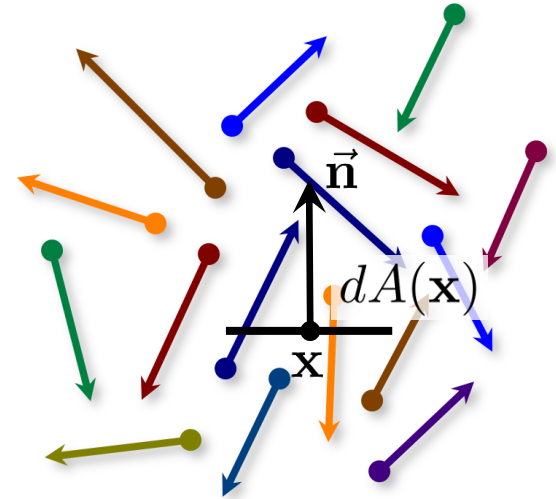
- Irradiance

- flux per unit area *arriving* at a surface = area density of flux

$$E(\mathbf{x}) = \frac{d\Phi(A)}{dA(\mathbf{x})} \left[ \frac{W}{m^2} \right]$$

- example:

- number of photons hitting a small patch of a wall per second, divided by the size of the patch



# Radiometry

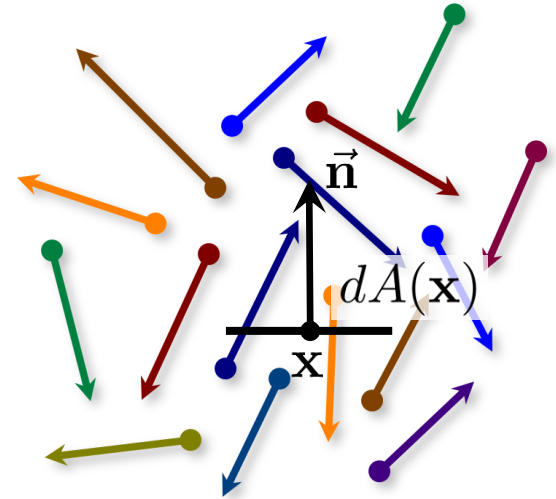
- Radiosity

- flux per unit area *leaving* a surface = area density of flux

$$B(\mathbf{x}) = \frac{d\Phi(A)}{dA(\mathbf{x})} \left[ \frac{W}{m^2} \right]$$

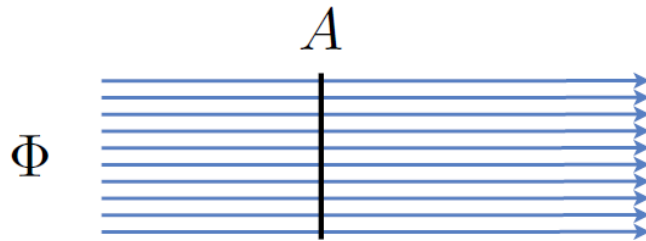
- example:

- number of photons reflecting off a small patch of a wall per second, divided by the size of the patch

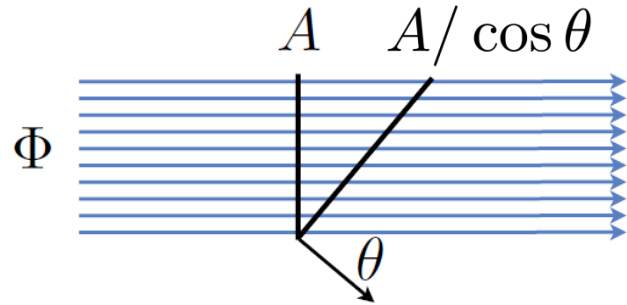


# Radiometry

- Irradiance
  - Lambert's Cosine Law



$$E = \frac{\Phi}{A}$$

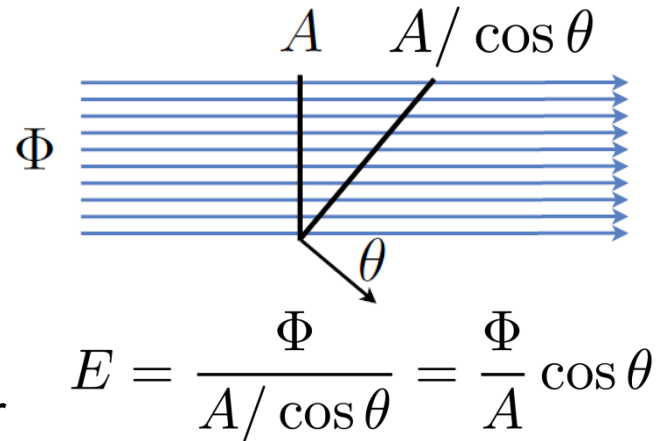
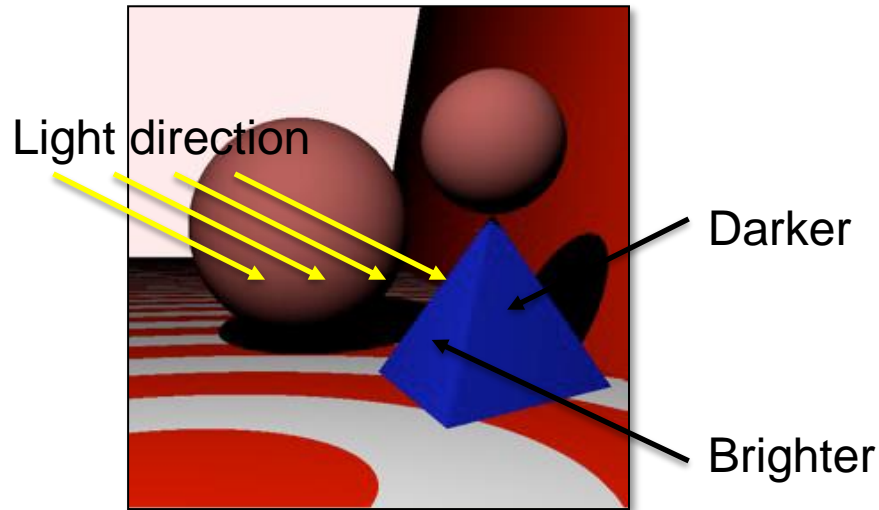


$$E = \frac{\Phi}{A / \cos \theta} = \frac{\Phi}{A} \cos \theta$$



# Radiometry

- Irradiance
  - Lambert's Cosine Law



# Radiometry

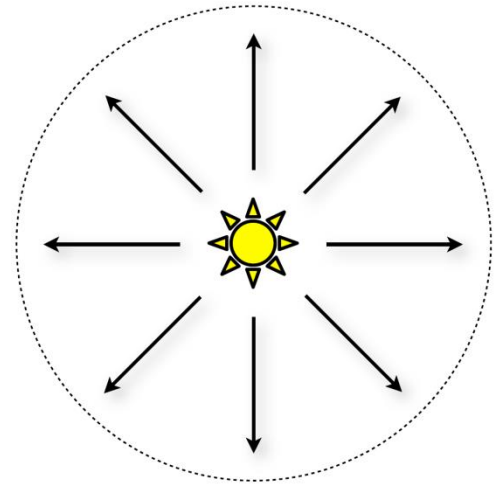
- Radiant intensity
  - Power (flux) per solid angle = directional density of flux

$$I(\vec{\omega}) = \frac{d\Phi}{d\vec{\omega}} \quad \left[ \frac{W}{sr} \right] \quad \Phi = \int_{S^2} I(\vec{\omega}) d\vec{\omega}$$

- example:

- power per unit solid angle emanating from a point source

$$\Phi = 4\pi I \quad \text{isotropic point source}$$



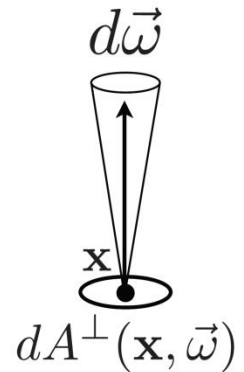
# Radiometry

- Radiance

- intensity per unit area = flux density per unit solid angle, per perpendicular unit area

$$L(\mathbf{x}, \vec{\omega}) = \frac{dI(\vec{\omega})}{dA(\mathbf{x})} = \frac{d^2\Phi(A)}{d\vec{\omega}dA^\perp(\mathbf{x}, \vec{\omega})} = \frac{d^2\Phi(A)}{d\vec{\omega}dA(\mathbf{x}) \cos \theta} \quad \left[ \frac{W}{m^2 sr} \right]$$

- most fundamental for raytracing
- remains constant along a ray



# Radiometry

- Other radiometric quantities can be expressed in terms of radiance

- Irradiance:  $L(\mathbf{x}, \vec{\omega}) = \frac{d^2\Phi(A)}{\cos\theta dA(\mathbf{x})d\vec{\omega}}$      $E(\mathbf{x}) = \frac{d\Phi(A)}{dA(\mathbf{x})}$

$$L(\mathbf{x}, \vec{\omega}) = \frac{dE(\mathbf{x})}{\cos\theta d\vec{\omega}}$$

$$L(\mathbf{x}, \vec{\omega}) \cos\theta d\vec{\omega} = dE(\mathbf{x})$$

$$\int_{H^2} L(\mathbf{x}, \vec{\omega}) \cos\theta d\vec{\omega} = E(\mathbf{x})$$

- Integrate radiance over the hemisphere
  - Same for radiosity

# Radiometry

- Other radiometric quantities can be expressed in terms of radiance

- Flux:

$$E(\mathbf{x}) = \frac{d\Phi(A)}{dA(\mathbf{x})}$$

$$\int_A E(\mathbf{x}) dA(\mathbf{x}) = \Phi(A)$$

$$\int_A \int_{H^2} L(\mathbf{x}, \vec{\omega}) \cos \theta d\vec{\omega} dA(\mathbf{x}) = \Phi(A)$$

- Integrate irradiance over area
    - Integrate radiance over hemisphere and area

# Radiometry

- Basic quantities

Flux

$$\Phi(A) \quad \left[ \frac{J}{s} = W \right]$$

Irradiance

$$E(\mathbf{x}) = \frac{d\Phi(A)}{dA(\mathbf{x})} \quad \left[ \frac{W}{m^2} \right]$$

Radiosity

$$B(\mathbf{x}) = \frac{d\Phi(A)}{dA(\mathbf{x})} \quad \left[ \frac{W}{m^2} \right]$$

Intensity

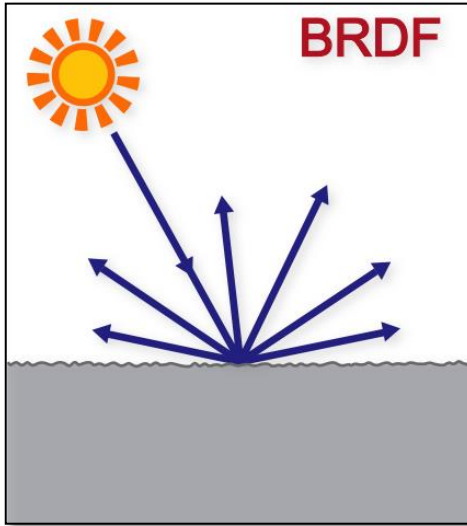
$$I(\vec{\omega}) = \frac{d\Phi}{d\vec{\omega}} \quad \left[ \frac{W}{sr} \right]$$

Radiance

$$L(\mathbf{x}, \vec{\omega}) = \frac{d^2\Phi(A)}{\cos\theta dA(\mathbf{x})d\vec{\omega}} \quad \left[ \frac{W}{m^2 sr} \right]$$

# Reflection Models

- **Bidirectional Reflectance Distribution Function (BRDF)**

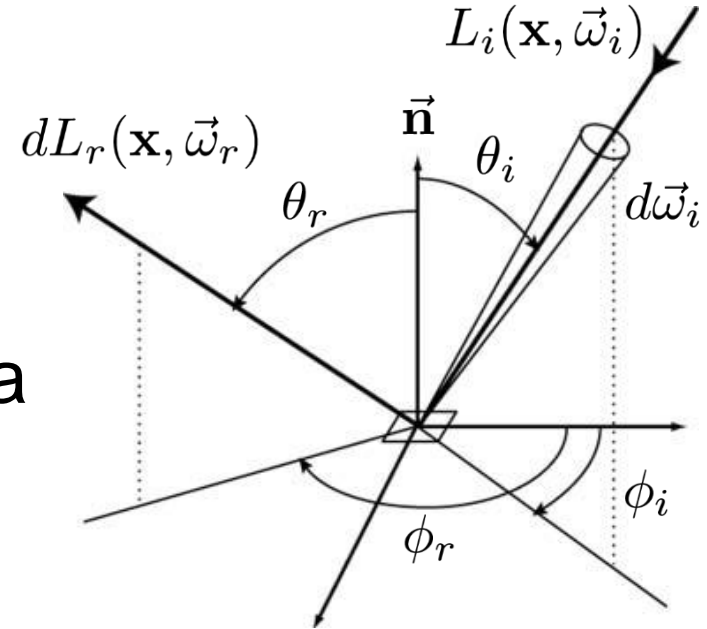


# BRDF

- **Bidirectional Reflectance Distribution Function**

$$\begin{aligned} f_r(\mathbf{x}, \vec{\omega}_i, \vec{\omega}_r) &= \frac{dL_r(\mathbf{x}, \vec{\omega}_r)}{dE_i(\mathbf{x}, \vec{\omega}_i)} \\ &= \frac{dL_r(\mathbf{x}, \vec{\omega}_r)}{L_i(\mathbf{x}, \vec{\omega}_i) \cos \theta_i d\vec{\omega}_i} \quad [1/sr] \end{aligned}$$

- Differential irradiance due to a cone of directions around  $\vec{\omega}_i$





# Reflection Equation

- The BRDF provides a relation between incident radiance and differential reflected radiance
- From this we can derive the **Reflection Equation**

$$f_r(\mathbf{x}, \vec{\omega}_i, \vec{\omega}_r) = \frac{dL_r(\mathbf{x}, \vec{\omega}_r)}{L_i(\mathbf{x}, \vec{\omega}_i) \cos \theta_i d\vec{\omega}_i}$$

$$f_r(\mathbf{x}, \vec{\omega}_i, \vec{\omega}_r) L_i(\mathbf{x}, \vec{\omega}_i) \cos \theta_i = \frac{dL_r(\mathbf{x}, \vec{\omega}_r)}{d\vec{\omega}_i}$$

$$\int_{H^2} f_r(\mathbf{x}, \vec{\omega}_i, \vec{\omega}_r) L_i(\mathbf{x}, \vec{\omega}_i) \cos \theta_i d\vec{\omega}_i = L_r(\mathbf{x}, \vec{\omega}_r)$$

# Reflection Equation

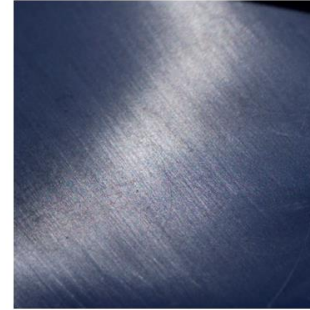
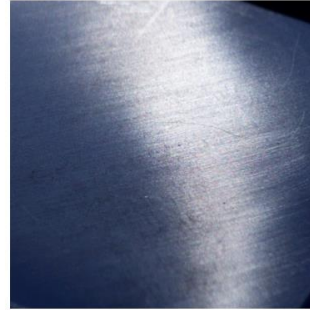
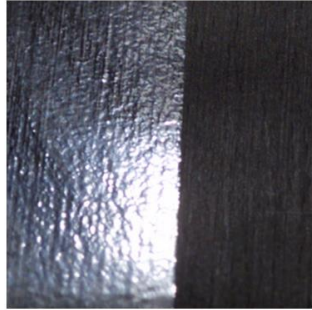
- The Reflection Equation describes a *local illumination* model
  - reflected radiance due to incident illumination from all directions

$$L_r(\mathbf{x}, \vec{\omega}_r) = \int_{H^2} f_r(\mathbf{x}, \vec{\omega}_i, \vec{\omega}_r) L_i(\mathbf{x}, \vec{\omega}_i) \cos \theta_i d\vec{\omega}_i$$

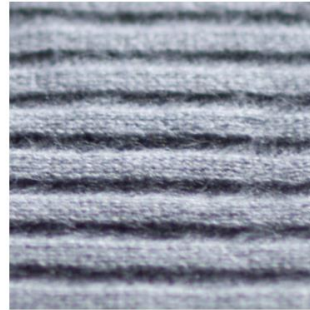
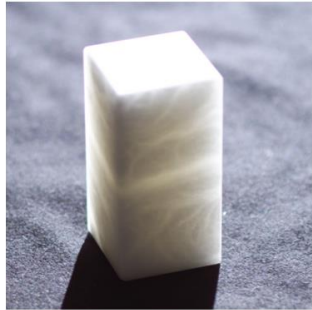
# Complex Reflections

anisotropic reflections

layered  
materials

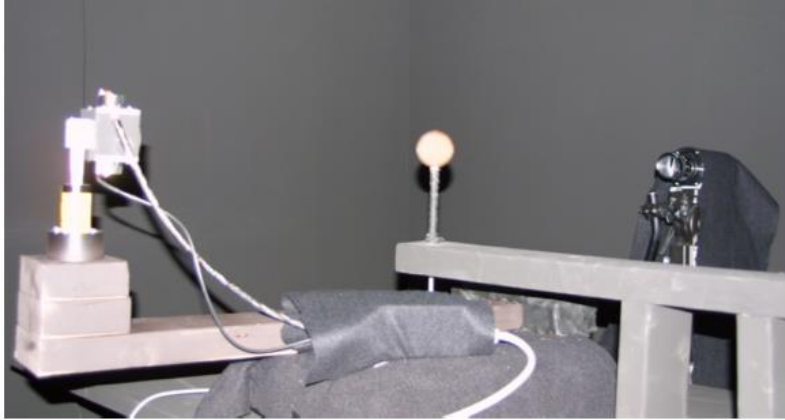


subsurface  
scattering

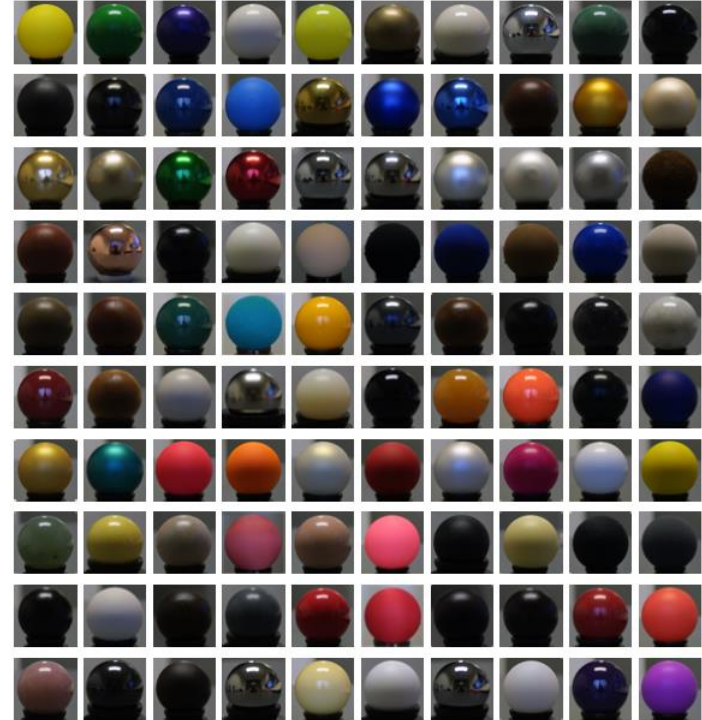


volumetric structures

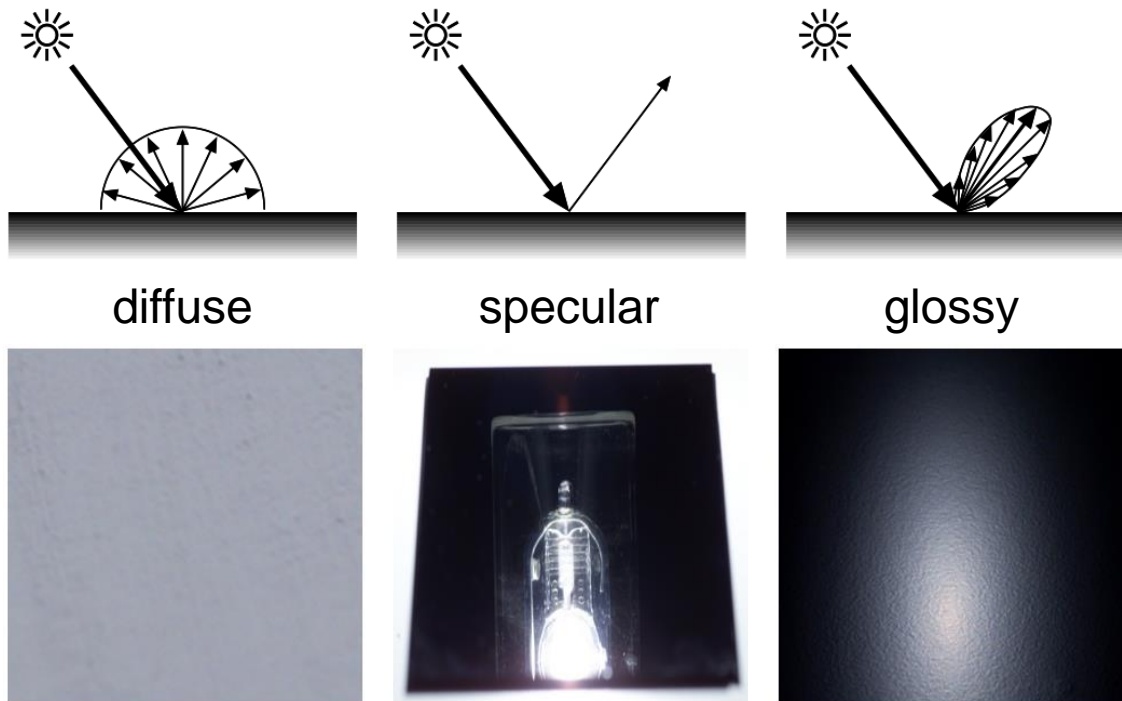
# Measuring BRDFs



Matusik et al.: Efficient Isotropic BRDF Measurement, Eurographics Symposium on Rendering 2003



# Simpler Reflections



Hendrik Lensch, Efficient Image-Based Appearance Acquisition of Real-World Objects, Ph.D. thesis, 2004

# Diffuse Reflection

- For diffuse reflection, the BRDF is a constant:

$$L_r(\mathbf{x}, \vec{\omega}_r) = \int_{H^2} f_r(\mathbf{x}, \vec{\omega}_i, \vec{\omega}_r) L_i(\mathbf{x}, \vec{\omega}_i) \cos \theta_i d\vec{\omega}_i$$

$$L_r(\mathbf{x}) = f_r \int_{H^2} L_i(\mathbf{x}, \vec{\omega}_i) \cos \theta_i d\vec{\omega}_i$$

$$L_r(\mathbf{x}) = f_r E_i(\mathbf{x})$$

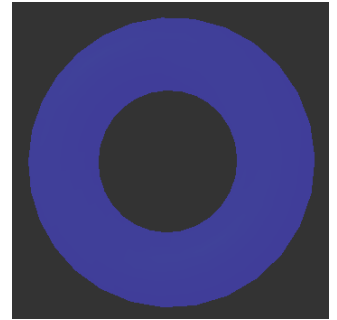
# Simple Models

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- Exact computation too slow
- OpenGL uses simplified reflection models
- Phong illumination

# Ambient Light

- Scattered by environment
- Coming from all directions
- Reflection independent of
  - Camera position
  - Light position (no light position)
  - Surface orientation
- Reflected intensity:  $I = I_a k_a$



light source

material parameter

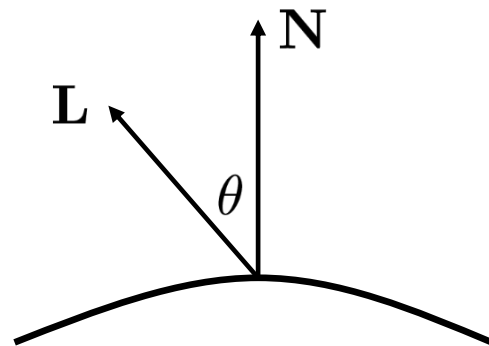
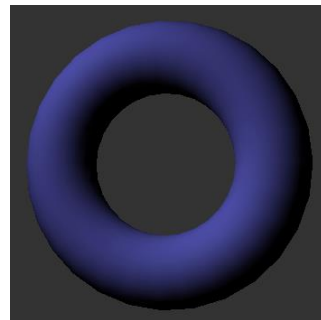


# Diffuse Reflection

- Directed light  $I_p$
- Reflection dependent on
  - orientation of surface
  - light source position
- Independent of
  - camera position (reflected equally in all directions)
- Reflected intensity:

$$I = I_p k_d \cos \theta$$

$$I = I_p k_d (\mathbf{N} \cdot \mathbf{L})$$

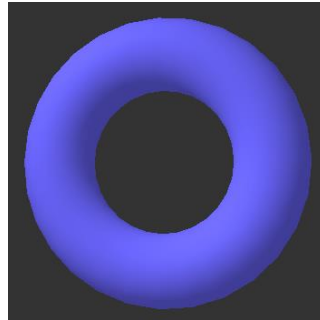


ve dot product or

# A Simple Model

- Sum up ambient light and diffuse reflection:

$$I = I_a k_a + I_p k_d (\mathbf{N} \cdot \mathbf{L})$$



# Attenuation

- Quadratic attenuation due to spatial radiation

$$f_{att} = \frac{1}{d_L^2}$$

- A model often used in Graphics (OpenGL)

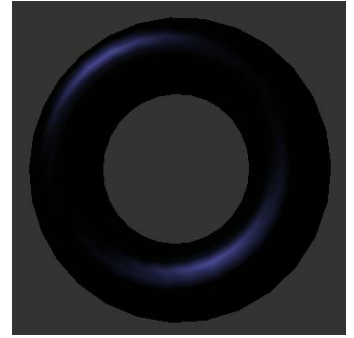
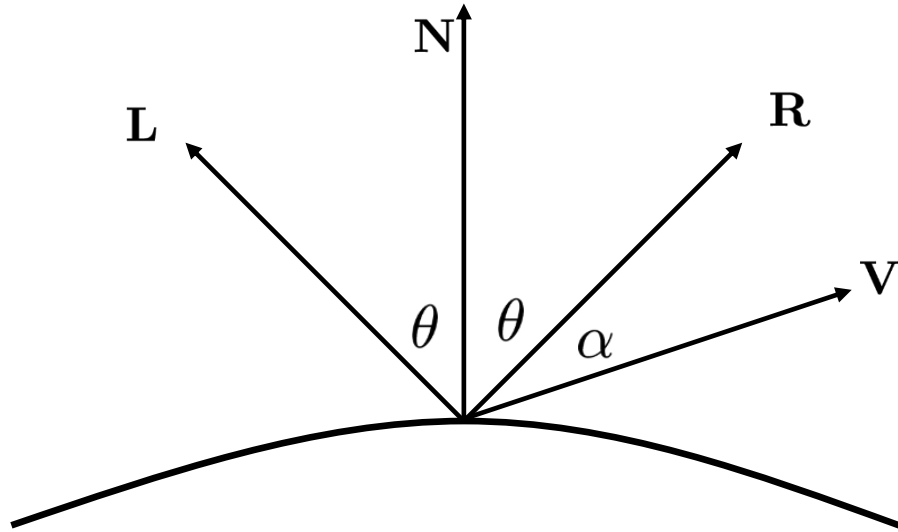
$$f_{att} = \min \left( \frac{1}{c_1 + c_2 d_L + c_3 d_L^2}, 1 \right)$$

- Include attenuation

$$I = I_a k_a + f_{att} I_p k_d (\mathbf{N} \cdot \mathbf{L})$$

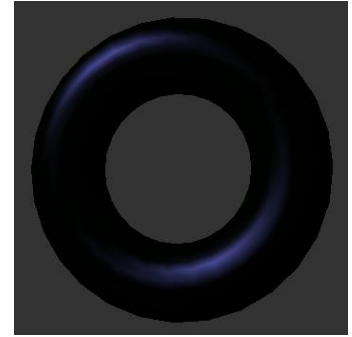
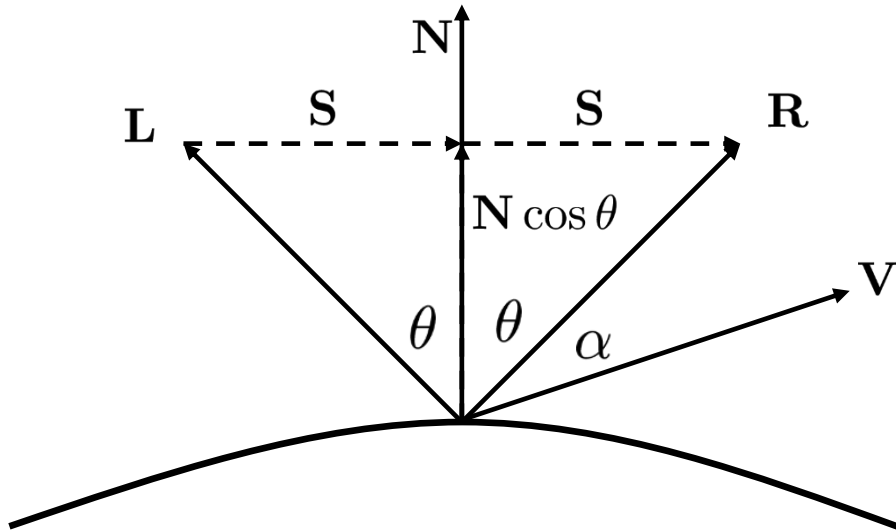
# Specular Reflection

- Depends on the angle between the reflection and viewing ray



# Specular Reflection

- Compute by simple linear algebra

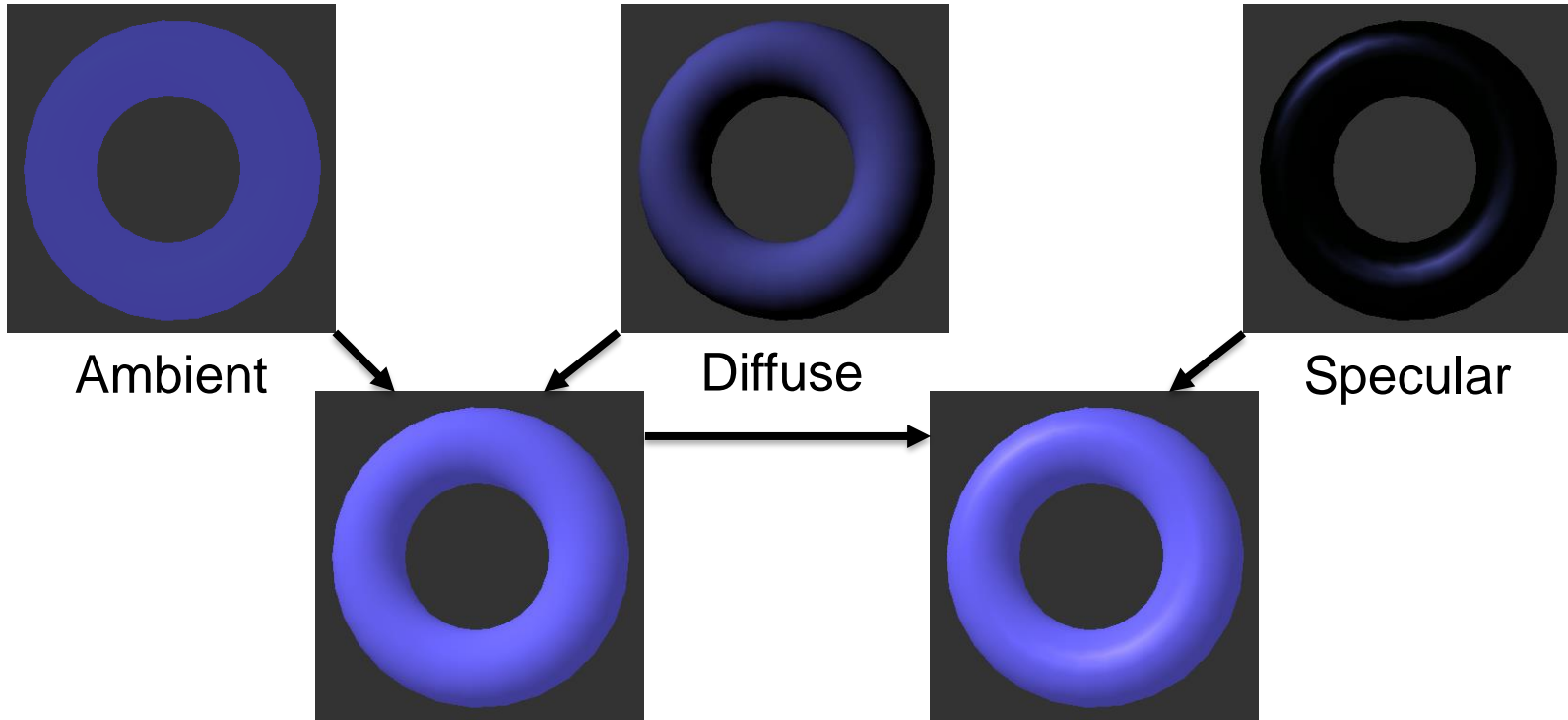


$$\mathbf{R} = \mathbf{N} \cos \theta + \mathbf{S}$$

$$\mathbf{R} = 2\mathbf{N} \cos \theta - \mathbf{L} = 2\mathbf{N}(\mathbf{N} \cdot \mathbf{L}) - \mathbf{L}$$

$$\cos \alpha = \mathbf{R} \cdot \mathbf{V} = (2\mathbf{N}(\mathbf{N} \cdot \mathbf{L}) - \mathbf{L}) \cdot \mathbf{V}$$

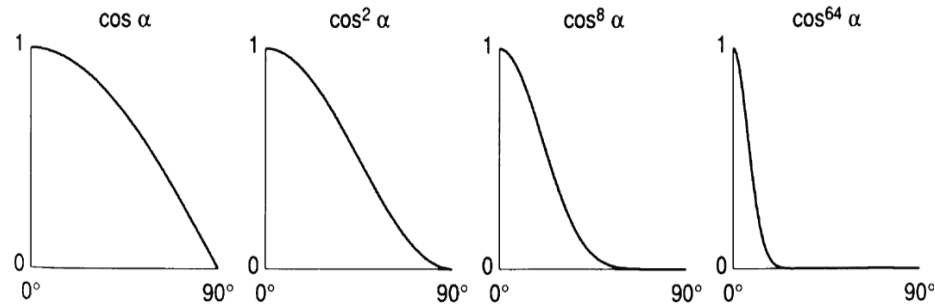
# Ambient + Diffuse + Specular



# Phong Illumination Model

- Approximates specular reflection by cosine powers

$$I_\lambda = I_{a_\lambda} k_a O_{d_\lambda} + f_{att} I_{p_\lambda} [k_d O_{d_\lambda} (\mathbf{N} \cdot \mathbf{L}) + k_s (\mathbf{R} \cdot \mathbf{V})^n]$$



# Extensions

- Specular colors

$$I_\lambda = I_{a_\lambda} k_a + f_{att} I_{p_\lambda} [k_d (\mathbf{N} \cdot \mathbf{L}) + k_s (\mathbf{R} \cdot \mathbf{V})^n]$$

Material dependent constants





# Extensions

- Specular colors

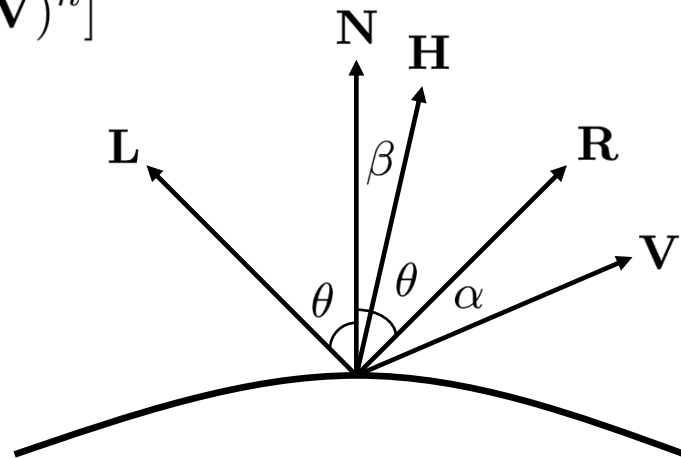
$$I_\lambda = I_{a_\lambda} k_a + f_{att} I_{p_\lambda} [k_d(\mathbf{N} \cdot \mathbf{L}) + k_s(\mathbf{R} \cdot \mathbf{V})^n]$$

- Halfway vector (faster)

$$\cos^n \beta = (\mathbf{N} \cdot \mathbf{H})^n \quad \mathbf{H} = \frac{\mathbf{L} + \mathbf{V}}{\|\mathbf{L} + \mathbf{V}\|}$$

- Multiple light sources

$$I_\lambda = I_{a_\lambda} k_a + \sum_{1 \leq i \leq m} f_{att_i} I_{p_{\lambda_i}} [k_d(\mathbf{N} \cdot \mathbf{L}_i) + k_s(\mathbf{R}_i \cdot \mathbf{V})^n]$$



# End

